

# Period-luminosity relations for Cepheid variables: from mid-infrared to multi-phase

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**Abstract** This paper discusses two aspects of current research on the Cepheid period-luminosity (P-L) relation: the derivation of mid-infrared (MIR) P-L relations and the investigation of multi-phase P-L relations.

The MIR P-L relations for Cepheids are important in the *James Webb Space Telescope* era for the distance scale issue, as the relations have potential to derive the Hubble constant within  $\sim 2\%$  accuracy—a critical constraint in precision cosmology. Consequently, we have derived the MIR P-L relations for Cepheids in the Large and Small Magellanic Clouds, using archival data from *Spitzer Space Telescope*. We also compared currently empirical P-L relations for Cepheids in the Magellanic Clouds to the synthetic MIR P-L relations derived from pulsational models.

For the study of multi-phase P-L relations, we present convincing evidence that the Cepheid P-L relations in the Magellanic Clouds are highly dynamic quantities that vary significantly when considered as a function of pulsational phase. We found that there is a difference in P-L relations as a function of phase between the Cepheids in each of the

Clouds; the most likely cause for this is the metallicity difference between the two galaxies. We also investigated the dispersion of the multi-phase P-L relations, and found that the minimum dispersions do not differ significantly from the mean light P-L dispersion.

**Keywords** Stars: variables: Cepheids · Distance scale

## 1 Introduction

The period-luminosity (P-L, also known as Leavitt Law) relation for Cepheid variables is an important astrophysical tool. A calibrated P-L relation can serve as the first rung in the extragalactic distance scale ladder, which can be used to determine the Hubble constant (e.g., Freedman et al. 2001; Sandage et al. 2006; Riess et al. 2011, and reference therein). In the local Universe, the Cepheid P-L relation can be used to measure the distances to nearby galaxies and investigate the characteristics of our own Galaxy (e.g., Majaess et al. 2009; Pedicelli et al. 2009, and reference therein). Research on Cepheid P-L relations includes calibrating the relations (e.g., Fouqué et al. 2007, and reference therein), investigating the metallicity dependence (e.g., Romaniello et al. 2008 and reference therein) or universality of the P-L relations (e.g., Bono et al. 2010, and reference therein), and the study of non-linearity of these relations (e.g., Ngeow et al. 2009, and reference therein). These works mainly focused on mean light in the optical and near infrared (*JHK*) bands. In this paper, we discuss two aspects of current research in P-L relations: the extension of the P-L relations to mid-infrared (Sect. 2), and the investigation of P-L relations at various phases of the pulsation—the multi-phase P-L relations (Sect. 3).

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